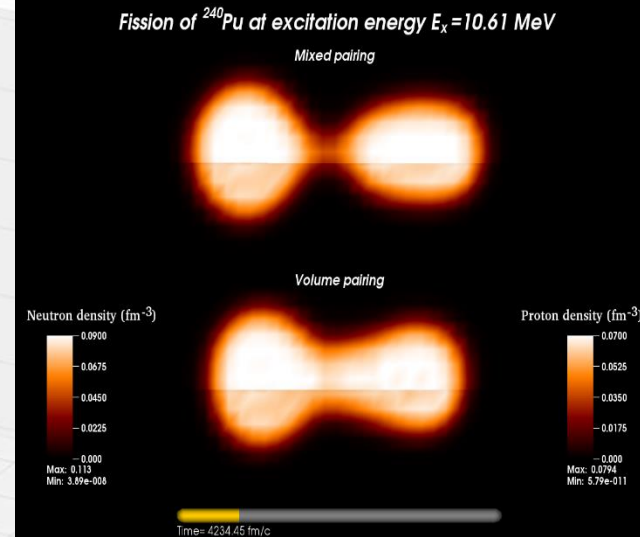
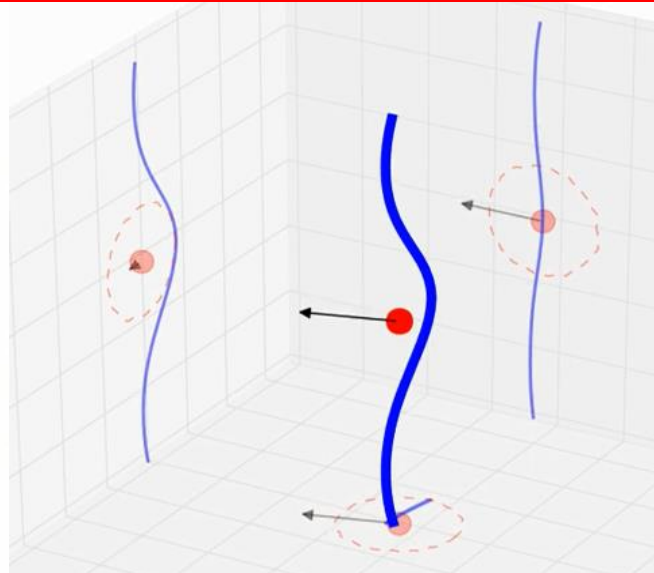
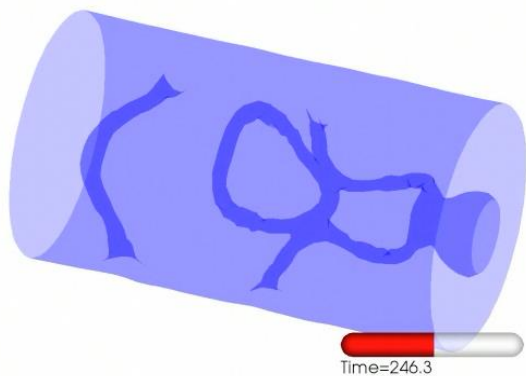


Selected applications of superfluid extension of TDDFT



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GOAL:

Description of superfluid dynamics far from equilibrium within the framework of Time Dependent Density Functional Theory (TDDFT).

We would like to describe the time evolution of (externally perturbed) spatially inhomogeneous, superfluid Fermi system and in particular such phenomena as:

- Vortex dynamics in ultracold Fermi gases and neutron matter.
- Vortex impurity interaction, vortex reconnections.
- Quantum turbulence.
- Atomic cloud collisions.
- Nuclear dynamics: large amplitude collective motion, induced nuclear fission, reactions, fusion, excitation of nuclei with gamma rays and neutrons.

Outline

- Basics of superfluid TDDFT in the local density approximation.
- Vortex dynamics in ultracold atomic gases.
- Vortex-impurity interaction in superfluid neutron matter.
- Nonlinear response of nuclear system: relativistic Coulomb excitation, induced fission.
- Effective mass of nuclear impurity in superfluid neutron matter.
- Solitonic excitations in nuclear reaction.

Runge Gross mapping

$$i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle = \hat{H} |\psi(t)\rangle, \quad |\psi_0\rangle = |\psi(t_0)\rangle$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \vec{j} = 0$$

$$\left. \begin{array}{l} \rho(\vec{r}, t) \\ |\psi(t_0)\rangle \end{array} \right\} \leftrightarrow e^{i\alpha(t)} |\psi(t)\rangle$$

Up to an arbitrary function $\alpha(t)$

and consequently the functional exists:

$$F[\psi_0, \rho] = \int_{t_0}^{t_1} \langle \psi[\rho] | \left(i\hbar \frac{\partial}{\partial t} - \hat{H} \right) | \psi[\rho] \rangle dt$$

Kohn-Sham approach


Suppose we are given the density of an interacting system.
There exists a unique noninteracting system with the same density.

Interacting system

$$i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle = (\hat{T} + \hat{V}(t) + \hat{W}) |\psi(t)\rangle$$

Noninteracting system

$$i\hbar \frac{\partial}{\partial t} |\varphi(t)\rangle = (\hat{T} + \hat{V}_{KS}(t)) |\varphi(t)\rangle$$


$$\rho(\vec{r}, t) = \langle \psi(t) | \hat{\rho}(\vec{r}) | \psi(t) \rangle = \langle \varphi(t) | \hat{\rho}(\vec{r}) | \varphi(t) \rangle$$

Hence the DFT approach is essentially exact.

However as always there is a price to pay:

- Kohn-Sham potential in principle depends on the past (memory).
Very little is known about the memory term and usually it is disregarded.
- Only one body observables can be reliably evaluated within standard DFT.

Pairing correlations in DFT

One may extend DFT to superfluid systems by defining the pairing field:

$$\Delta(\mathbf{r}\sigma, \mathbf{r}'\sigma') = -\frac{\delta E(\rho, \chi)}{\delta \chi^*(\mathbf{r}\sigma, \mathbf{r}'\sigma')}.$$

L. N. Oliveira, E. K. U. Gross, and W. Kohn, Phys. Rev. Lett. 60 2430 (1988).

O.-J. Wacker, R. Kümmel, E.K.U. Gross, Phys. Rev. Lett. 73, 2915 (1994).

and introducing anomalous density $\chi(\mathbf{r}\sigma, \mathbf{r}'\sigma') = \langle \hat{\psi}_{\sigma'}(\mathbf{r}') \hat{\psi}_{\sigma}(\mathbf{r}) \rangle$

However in the limit of the local field these quantities diverge unless one renormalizes the coupling constant:

$$\begin{aligned} \Delta(\mathbf{r}) &= g_{eff}(\mathbf{r}) \chi_c(\mathbf{r}) \\ \frac{1}{g_{eff}(\mathbf{r})} &= \frac{1}{g(\mathbf{r})} - \frac{mk_c(\mathbf{r})}{2\pi^2 \hbar^2} \left(1 - \frac{k_F(\mathbf{r})}{2k_c(\mathbf{r})} \ln \frac{k_c(\mathbf{r}) + k_F(\mathbf{r})}{k_c(\mathbf{r}) - k_F(\mathbf{r})} \right) \end{aligned}$$

which ensures that the term involving the kinetic and the pairing energy density is finite:

$$\frac{\tau_c(r)}{2m} - \Delta(r) \chi_c(r), \quad \tau_c(r) = \nabla \cdot \nabla' \rho_c(r, r') \Big|_{r=r'}$$

Bulgac, Yu, Phys. Rev. Lett. 88 (2002) 042504

Bulgac, Phys. Rev. C65 (2002) 051305

Formalism for Time Dependent Phenomena: TDSDA

Local density approximation (no memory terms)

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_{k\uparrow}(\mathbf{r}, t) \\ u_{k\downarrow}(\mathbf{r}, t) \\ v_{k\uparrow}(\mathbf{r}, t) \\ v_{k\downarrow}(\mathbf{r}, t) \end{pmatrix} = \begin{pmatrix} h_{\uparrow,\uparrow}(\mathbf{r}, t) & h_{\uparrow,\downarrow}(\mathbf{r}, t) & 0 & \Delta(\mathbf{r}, t) \\ h_{\downarrow,\uparrow}(\mathbf{r}, t) & h_{\downarrow,\downarrow}(\mathbf{r}, t) & -\Delta(\mathbf{r}, t) & 0 \\ 0 & -\Delta^*(\mathbf{r}, t) & -h_{\uparrow,\uparrow}^*(\mathbf{r}, t) & -h_{\uparrow,\downarrow}^*(\mathbf{r}, t) \\ \Delta^*(\mathbf{r}, t) & 0 & -h_{\downarrow,\uparrow}^*(\mathbf{r}, t) & -h_{\downarrow,\downarrow}^*(\mathbf{r}, t) \end{pmatrix} \begin{pmatrix} u_{k\uparrow}(\mathbf{r}, t) \\ u_{k\downarrow}(\mathbf{r}, t) \\ v_{k\uparrow}(\mathbf{r}, t) \\ v_{k\downarrow}(\mathbf{r}, t) \end{pmatrix}$$

Density functional contains normal densities, anomalous density (pairing) and currents:

$$E(t) = \int d^3r \left[\varepsilon(n(\vec{r}, t), \tau(\vec{r}, t), \nu(\vec{r}, t), \underline{\vec{j}}(\vec{r}, t)) + V_{ext}(\vec{r}, t)n(\vec{r}, t) + \dots \right]$$

- The system is placed on a large 3D spatial lattice.
- No symmetry restrictions
- Number of PDEs is of the order of the number of spatial lattice points

Current capabilities of the code:

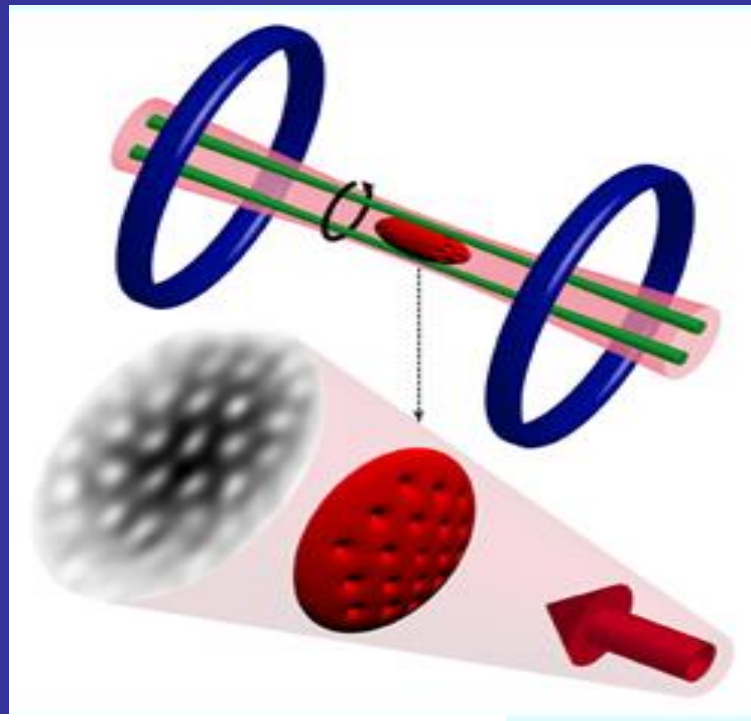
- volumes of the order of ($L = 80^3$) capable of simulating time evolution of 42000 neutrons at saturation density (natural application: neutron stars)
- For nuclear systems: capable of simulating up to times of the order of 10^{-19} s (a few million time steps)
- CPU vs GPU on Titan \approx 15 speed-up (likely an additional factor of 4 possible)

Eg. for 137062 two component wave functions:

CPU version (4096 nodes x 16 PEs) - 27.90 sec for 10 time steps

GPU version (4096 PEs + 4096GPU) - 1.84 sec for 10 time step

Vortex generation in ultracold Fermi gases

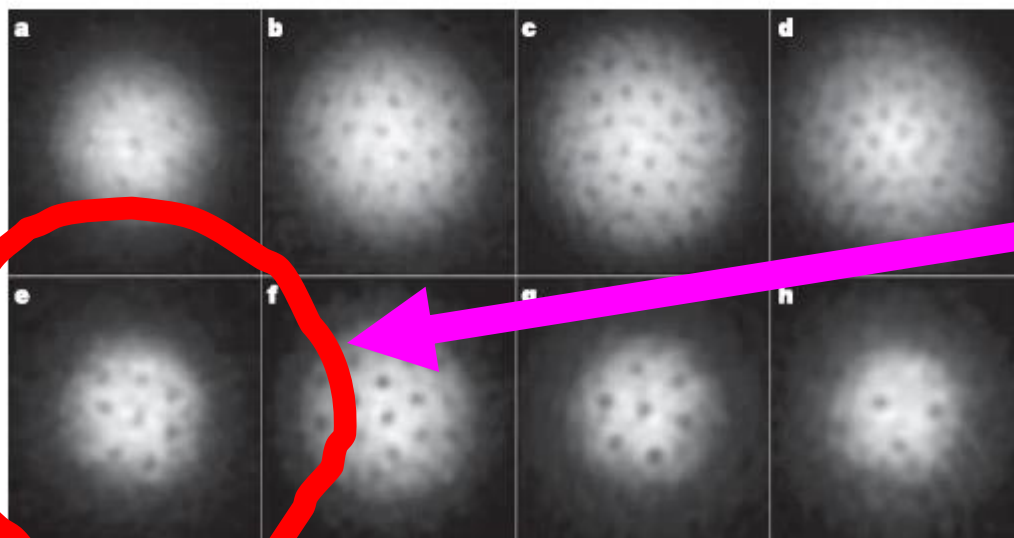


system of fermionic ${}^6\text{Li}$ atoms

Feshbach resonance:
 $B=834\text{G}$

BEC side:
 $a>0$

BCS side:
 $a<0$



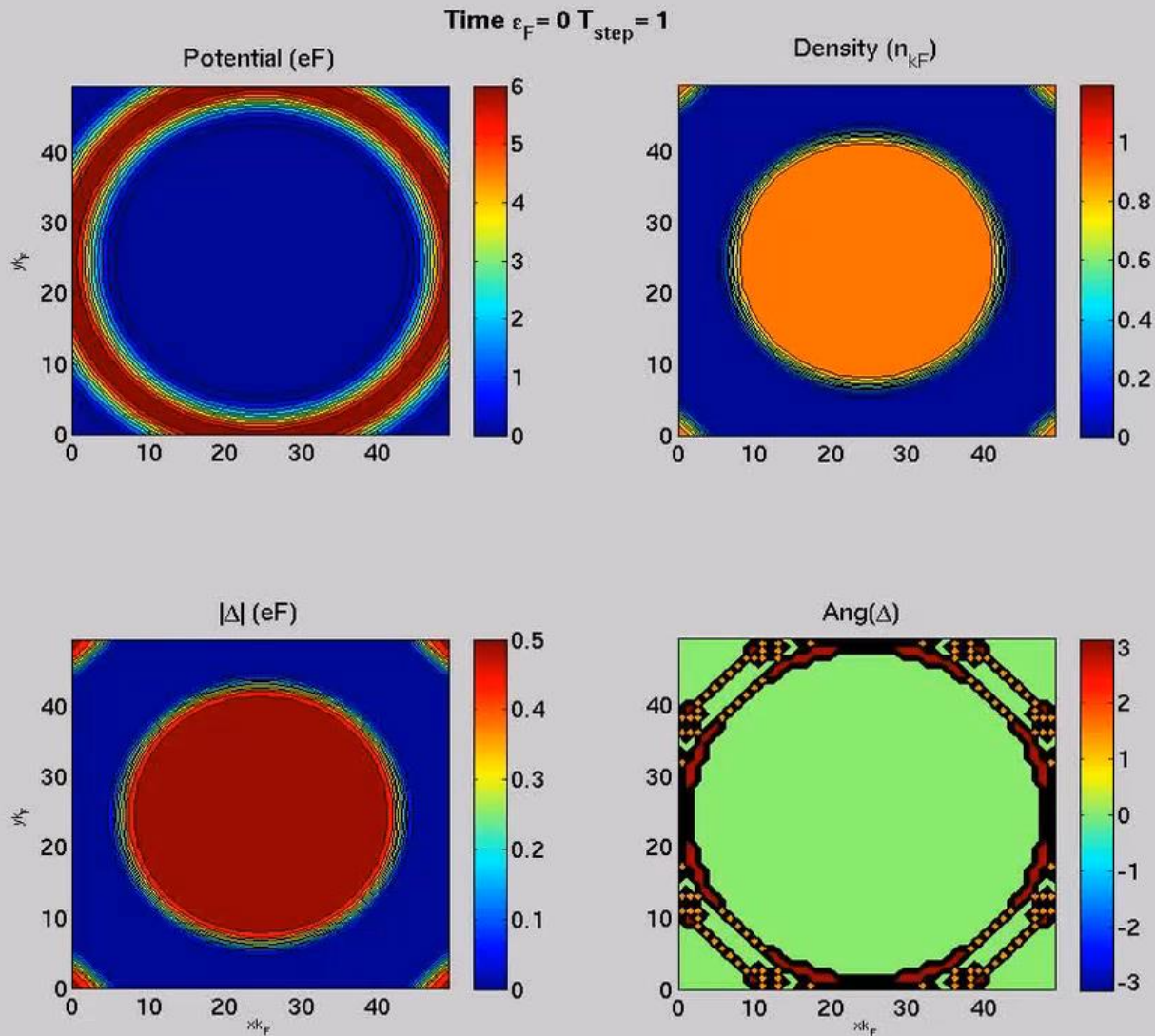
UNITARY REGIME

Figure 2 | Vortices in a strongly interacting Fermionic atoms on the BEC- and the BCS-side of the Feshbach resonance. At the given field, the cloud of lithium atoms was stirred for 300 ms (a) or 500 ms (b–h) followed by an equilibration time of 500 ms. After 2 ms of ballistic expansion, the

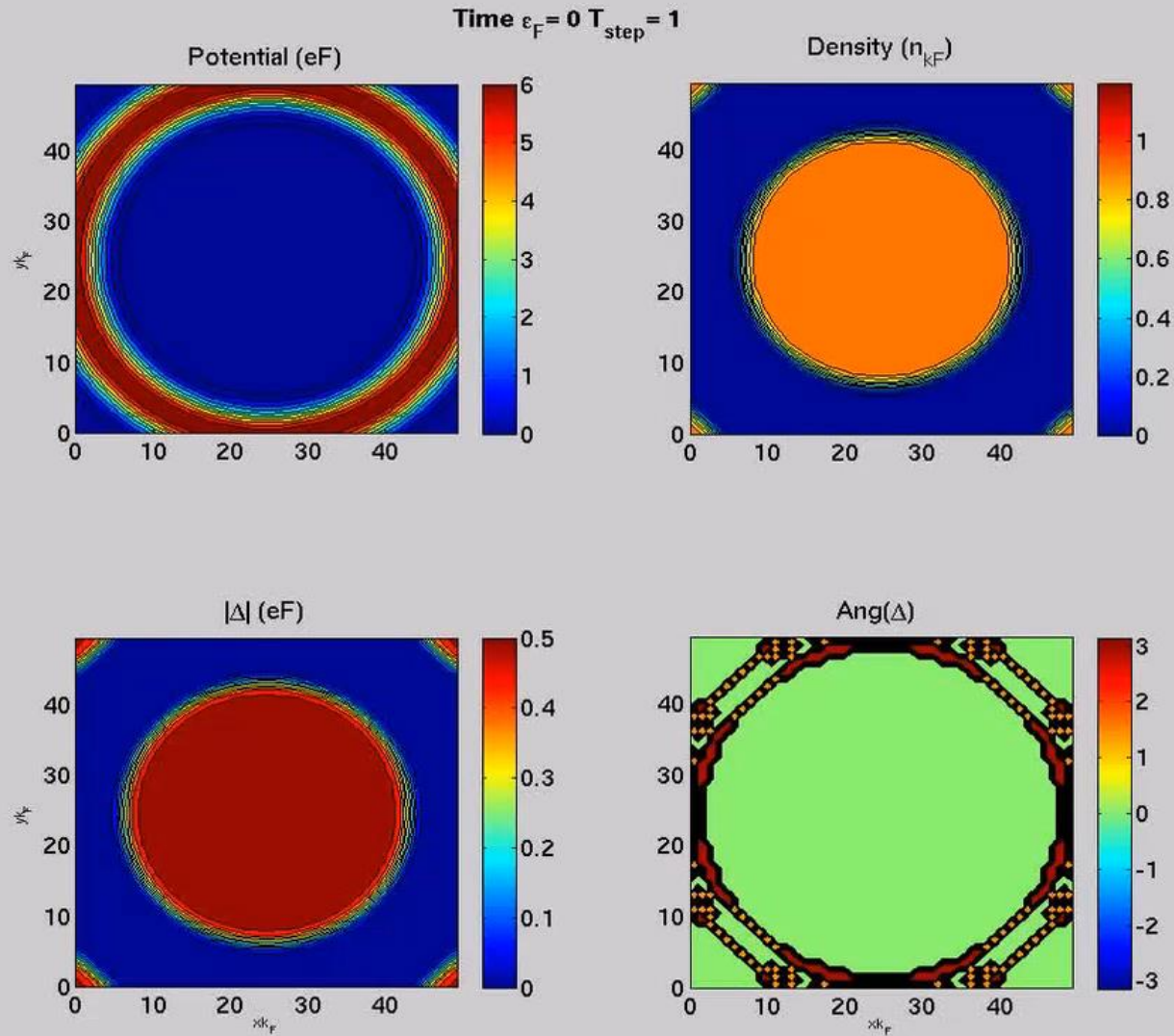
magnetic field was ramped to 735 G for imaging (see Methods). The magnetic fields were 740 G (a), 766 G (b), 792 G (c), 843 G (f), 853 G (g) and 863 G (h). The field of view was $880\ \mu\text{m} \times 880\ \mu\text{m}$.

M.W. Zwierlein *et al.*,
Nature, 435, 1047 (2005)

Stirring the atomic cloud with stirring velocity **lower** than the critical velocity



Stirring the atomic cloud with stirring velocity **exceeding** the critical velocity



Vortex reconnections

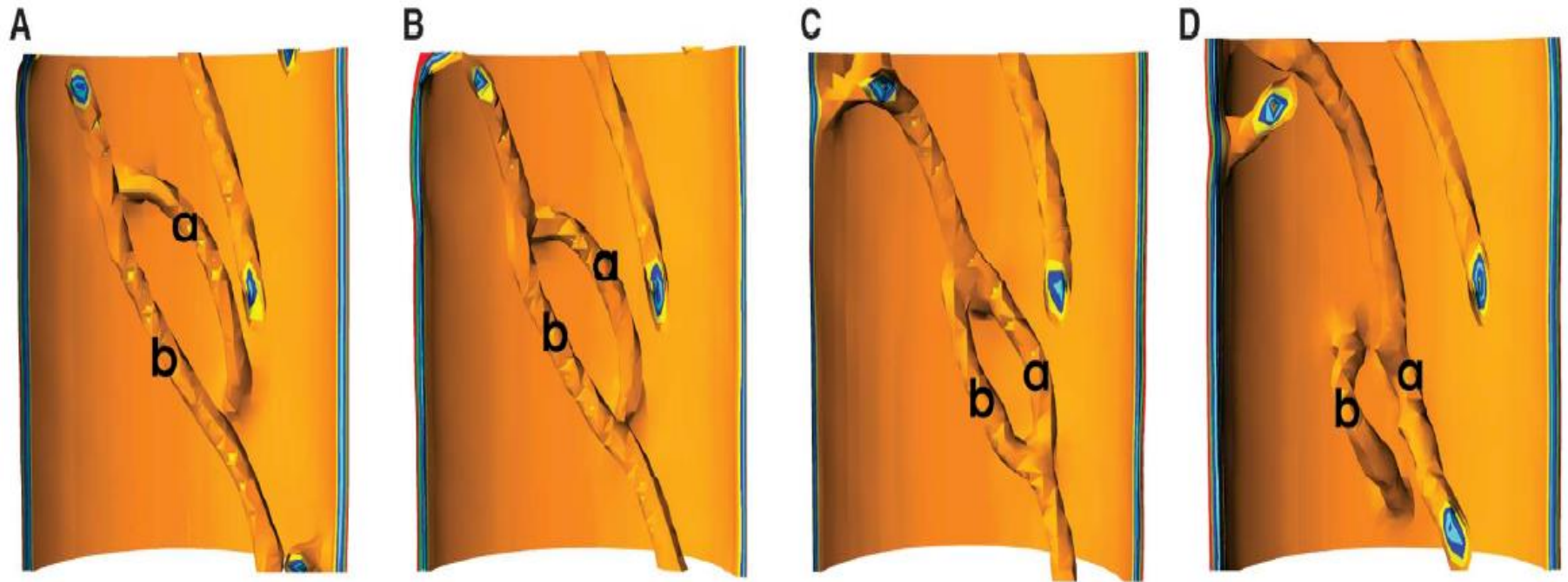
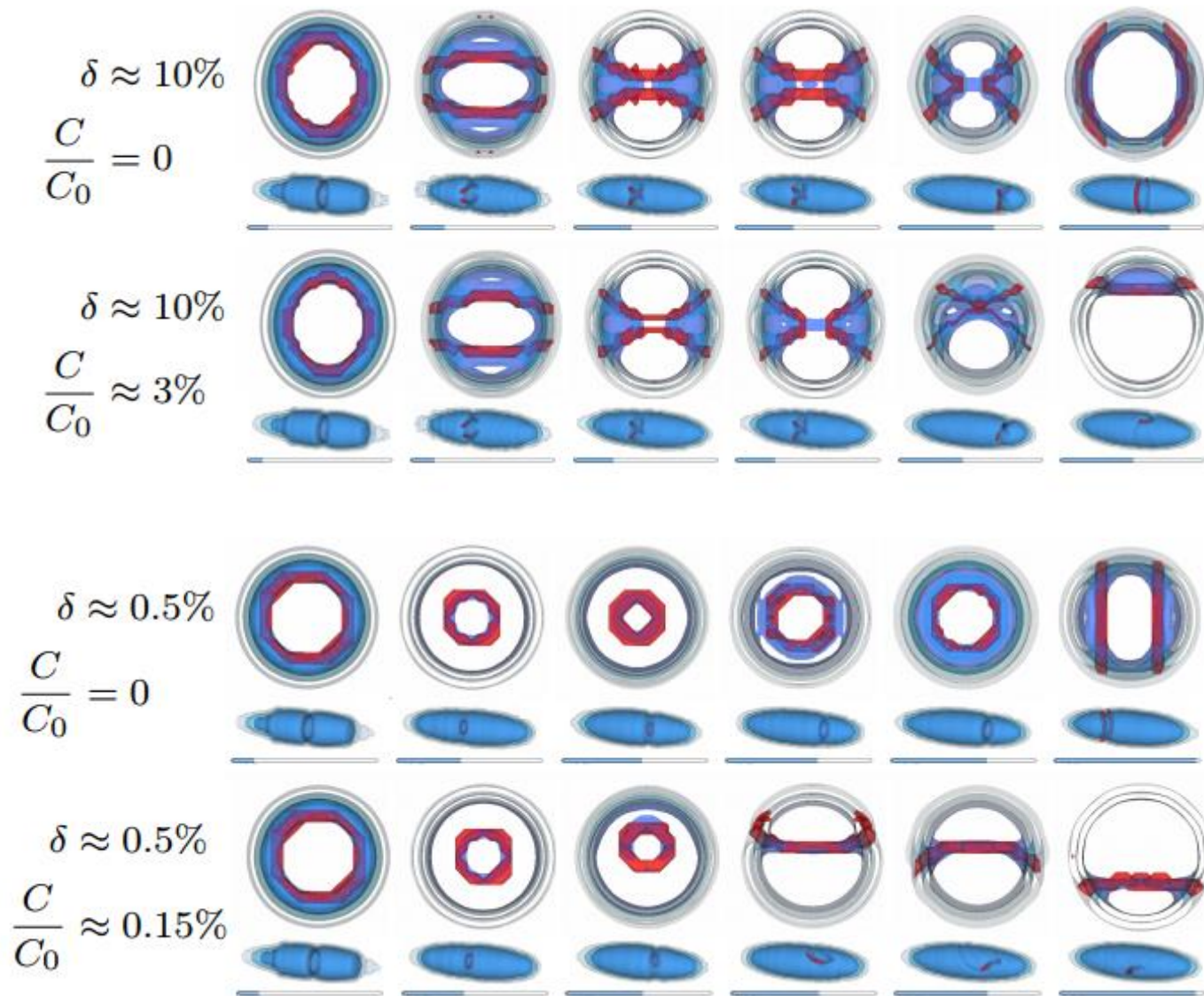


Fig. 3. (A to D) Two vortex lines approach each other, connect at two points, form a ring and exchange between them a portion of the vortex line, and subsequently separate. Segment (a), which initially belonged to the vortex line attached to the wall, is transferred to the long vortex line (b) after reconnection and vice versa.

Vortex reconnections are important for the energy dissipation mechanism in quantum turbulence.

TDSLDA can describe these processes as well as the energy transfer between collective and single particle degrees of freedom (which is a problem for simplified treatments based e.g. on Gross-Pitaevskii equation)



Moreover with TDDFT we can reproduce the sequence of topological excitations observed experimentally (M.H.J. Ku et al. Phys. Rev. Lett. 113, 065301 (2014)).

Vortex dynamics and vortex-impurity interaction

The effective equations of motion for the vortex dynamics (per unit length of the vortex):

$$M_{\text{vor}} \frac{d^2 \vec{r}}{dt^2} = \vec{F}_M + \vec{F}_D + \vec{F}_{\text{vor-impurity}}$$

$\vec{F}_M = \rho_s \vec{\Gamma} \times \left(\frac{d\vec{r}}{dt} - \vec{v}_s \right)$ - Magnus force; $\vec{\Gamma}$ - local vorticity;

$\frac{d\vec{r}}{dt}$ - local vortex velocity, ρ_s - superfluid density, \vec{v}_s - superfluid velocity

\vec{F}_D - frictional force (negligible at small T)

$\vec{F}_{\text{vor-impurity}}$ - vortex-impurity force

To date the impurity-vortex interaction has been extracted from static calculations (Ginzburg-Landau, local density, HFB) with several severe approximations:

- Vortex is always straight
- Nucleus is spherical
- Only very symmetric configurations are considered:
 - nucleus on vortex
 - vortex inbetween two nuclei (interstitial configuration)
 - nucleus at infinity

M.A. Alpar et al. *Astrophys.J.*213,527(1977);276,325(1984)
R.I. Epstein, G.Baym, *Astrophys.J.*328,680(1988)
R.K.Link,R.I.Epstein,*Astrophys.J.*373,592(1991)
P.Pizzochero et al. *PRL* 79,3347(1997)
R.Brogia et al.*PRD*50,4781(1994)
M.Baldo et al. *Nucl.Phys.*515,409(1990)
P.Donati,P.Pizzochero, *PRL*90,211101(2003);
*PLB*640,74(2006);*Nucl.Phys.A*742,363(2004)
P.Avogadro et al. *PRC*75,012805(2007);*NPA*811,378(2008)

From these assumptions the average force can be deduced and the energetically favorable configuration may be determined.

We use approach based on TDDFT which allow to extract the force from dynamics.

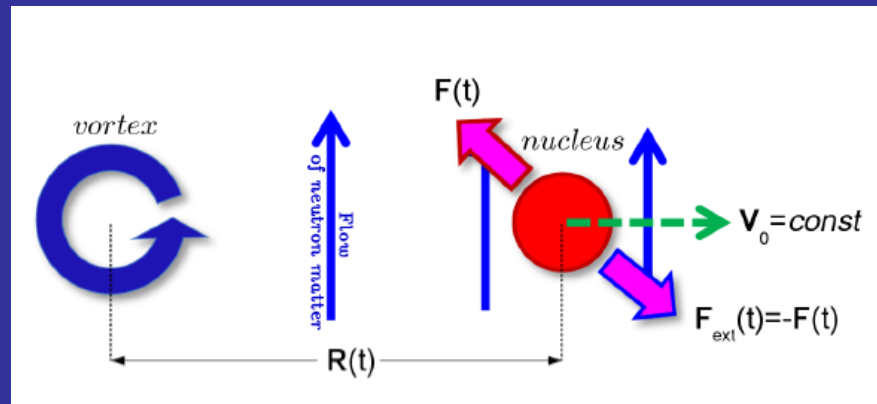
It has the following advantages:

- We can extract the force at various nonsymmetric configurations
- All degrees of freedom are treated on the same footing and in particular those associated with the vortex (bending) and nucleus (deformations) are taken into account
- One can get a better insight into the dynamics of the vortex-impurity system at various energy scales.

The procedure consists of dragging protons through the neutron medium with the vortex.

Such an approach has been shown to give the same force as extracted from static configurations (Bulgac,Forbes,Sharma, PRL 110,241102(2013)).

It is numerically much cheaper than searching for stationary solutions at various vortex-impurity configurations.



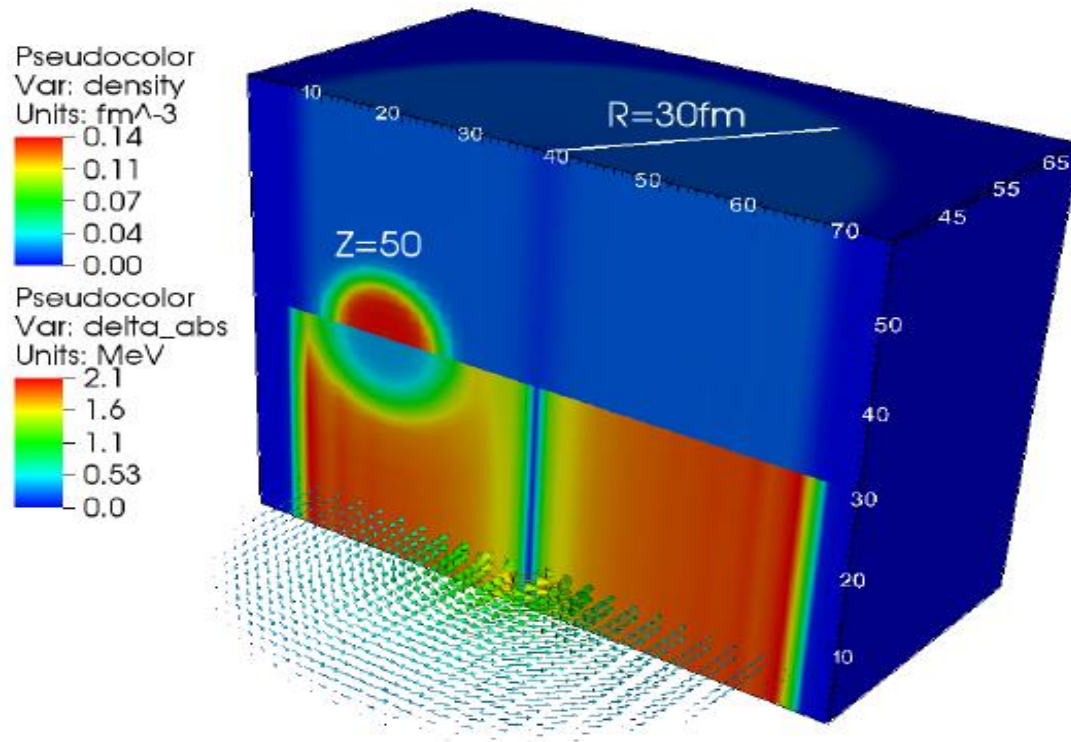
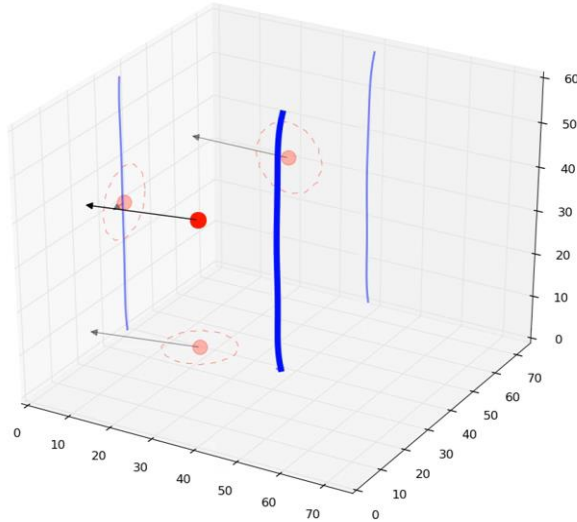


FIG. 5: (Color online) Example of initial unpinned configuration for $n = 0.014 \text{ fm}^{-3}$. The upper part of the box shows the total density distributions, while the lower part presents the absolute value of the neutron pairing potential Δ . The vanishing pairing field and the depletion of the density along the tube symmetry axis are due to the presence of a quantum vortex. Arrows in the bottom part of the figure show the circular flow of neutrons.

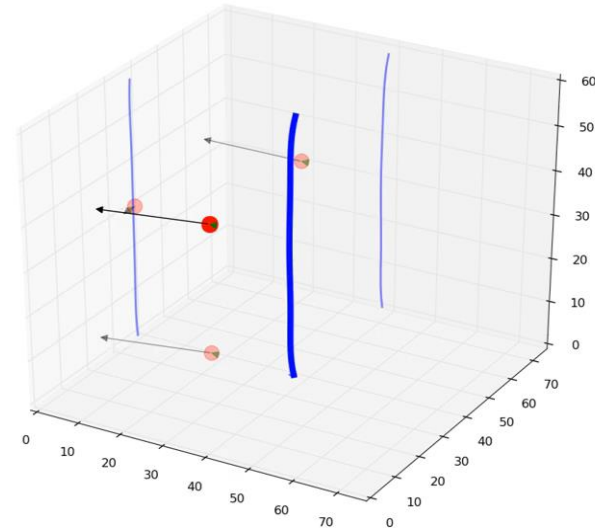
Vortex – impurity interaction

The external potential: $V_{EXT}(\vec{r}) = -\vec{F} \cdot \vec{r}$ keeps the nucleus moving along the straight line with a constant velocity below the critical velocity.

time= 0 fm/c
 F(19.1)= 2.08 MeV/fm
 Q= 28.0fm²



time= 11 fm/c
 F_m(19.1)= 2.08 MeV/fm
 F_t(19.1)= 0.01 MeV/fm



One can extract the total force and also the force exerted on each part of the vortex. Assuming that the force behaves asymptotically as $1/r^3$ (superfluid hydrodynamic estimate) one can extract the force per unit length by fitting the expression :

$$f(r) = \frac{\sum_{k=0}^n a_n r^n}{\sum_{k=0}^{n+3} b_n r^n}; f(r) \sim \frac{1}{r^3}; b_0 = 1$$

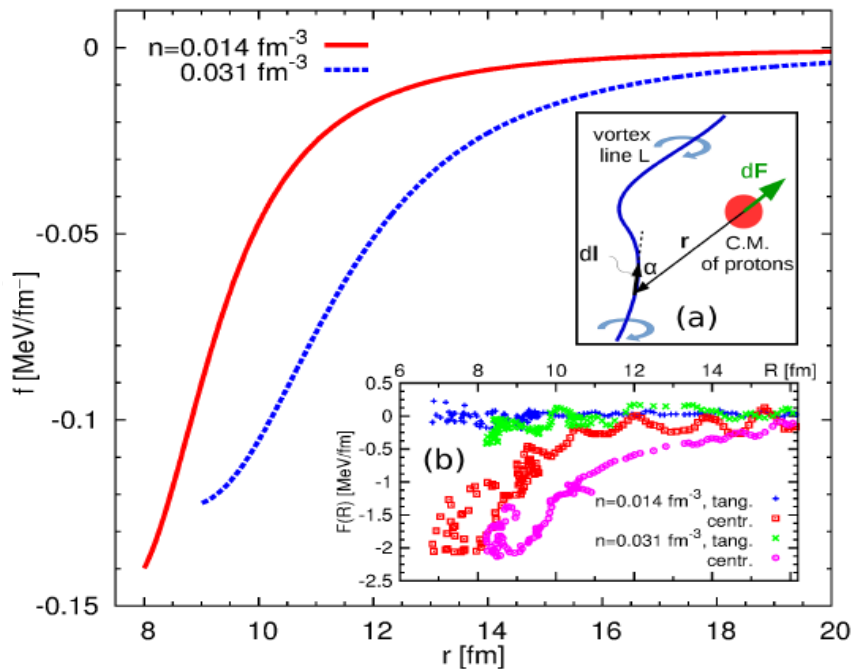


FIG. 3: (Color online) Extracted force per unit length $f(r)$ for different densities. Negative values means repulsive nature of the force. In inset (a) sketch explaining meaning of Eq. (4) is shown. Inset (b) shows the measured total force $F(R)$ as shown in Fig. 1 for different densities. The force has been decomposed into tangential and centripetal components with respect of temporal vortex position.

Vortex tension (upper limit): how much energy one needs to deform the vortex



Vortex-impurity repulsive force

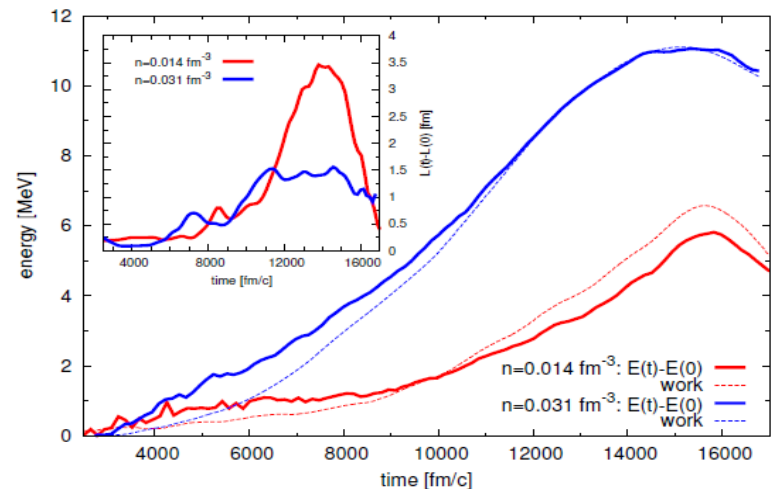


FIG. 8: (Color online) Excitation energy $E(t) - E(0)$ of the system as a function of time. By dashed line work performed by external force computed by formula $W(t) = \int_0^t \mathbf{F}_{\text{ext}}(t') \cdot \mathbf{v}(t') dt'$ is presented. In inset change of vortex length $L(t) - L(0)$ is shown as function of time.

$$T = 1.4 \text{ MeV} / \text{fm} \quad \text{for } n=0.014 \text{ fm}^{-3}$$

$$T = 7.3 \text{ MeV} / \text{fm} \quad \text{for } n=0.031 \text{ fm}^{-3}$$

Effective mass of a nucleus in superfluid neutron environment

Suppose we would like to evaluate an effective mass of a heavy particle immersed in a uniform Fermi bath.

Can one come up with the effective (classical) equation of motion of the type:

$$M_{\text{eff}} \frac{d^2 q}{dt^2} - F_D \left(\frac{dq}{dt}, \dots \right) = 0 \quad ?$$

In general it is a complicated task as the first and the second term may not be unambiguously separated.

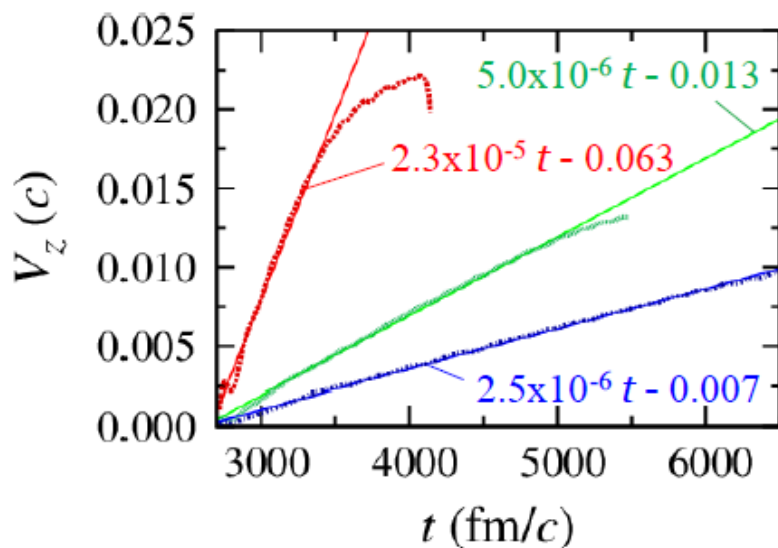
However for the superfluid system it can be done as for sufficiently slow motion (below the critical velocity) the second term may be neglected due to the presence of the pairing gap.

Effective mass of a nucleus immersed in superfluid neutron matter:

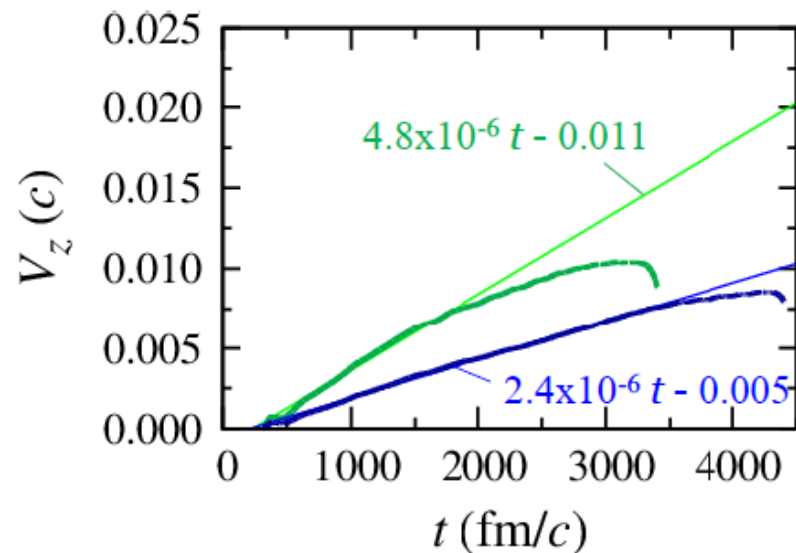
We apply the external potential for protons (50) of the form: $V_{EXT}(\vec{r}) = -\vec{F} \cdot \vec{r}$

and measure the C.M. velocity of the system

Lower density: $\rho_n \sim 0.016 \text{ fm}^{-3}$



Higher density: $\rho_n \sim 0.032 \text{ fm}^{-3}$



At low velocities: $v(t) \sim t$

The deviations for higher velocities indicate excitations of other modes

Dragging 50 protons the effective mass corresponds to dragging about 207 nucleons for lower density and about 228 nucleons for higher density.

Linear response regime

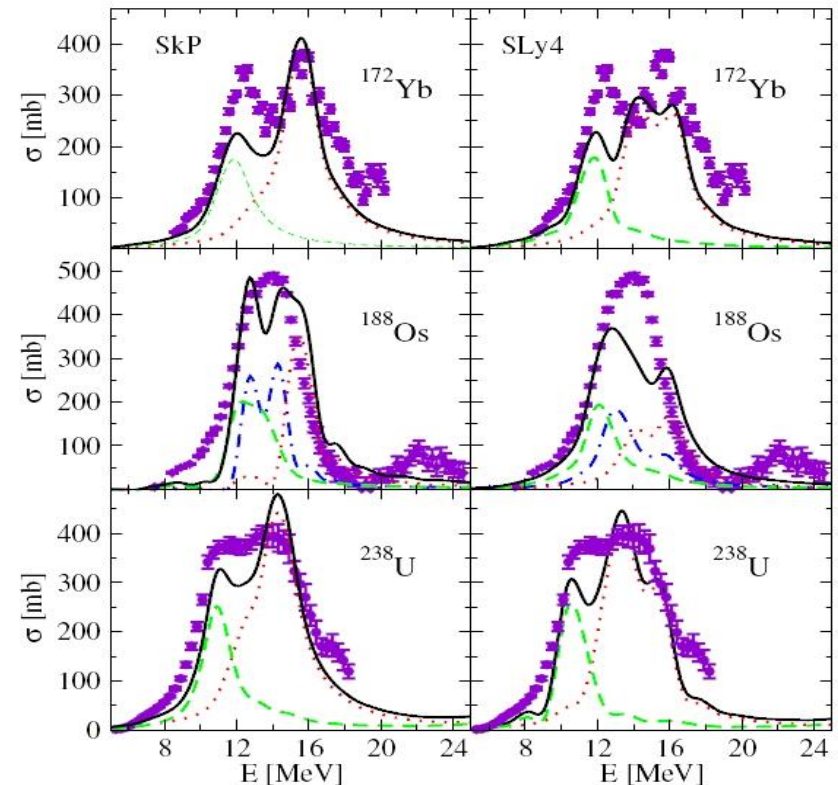
Photoabsorption cross section for heavy, deformed nuclei.

$$h_{\tau,\sigma\sigma}(\mathbf{r}, t) \Rightarrow h_{\tau,\sigma\sigma}(\mathbf{r}, t) + F_{\tau}(\mathbf{r})f(t) \quad F_{\tau}(\mathbf{r}) = N_{\tau} \sin(\mathbf{k} \cdot \mathbf{r}_{\tau})/|\mathbf{k}|$$

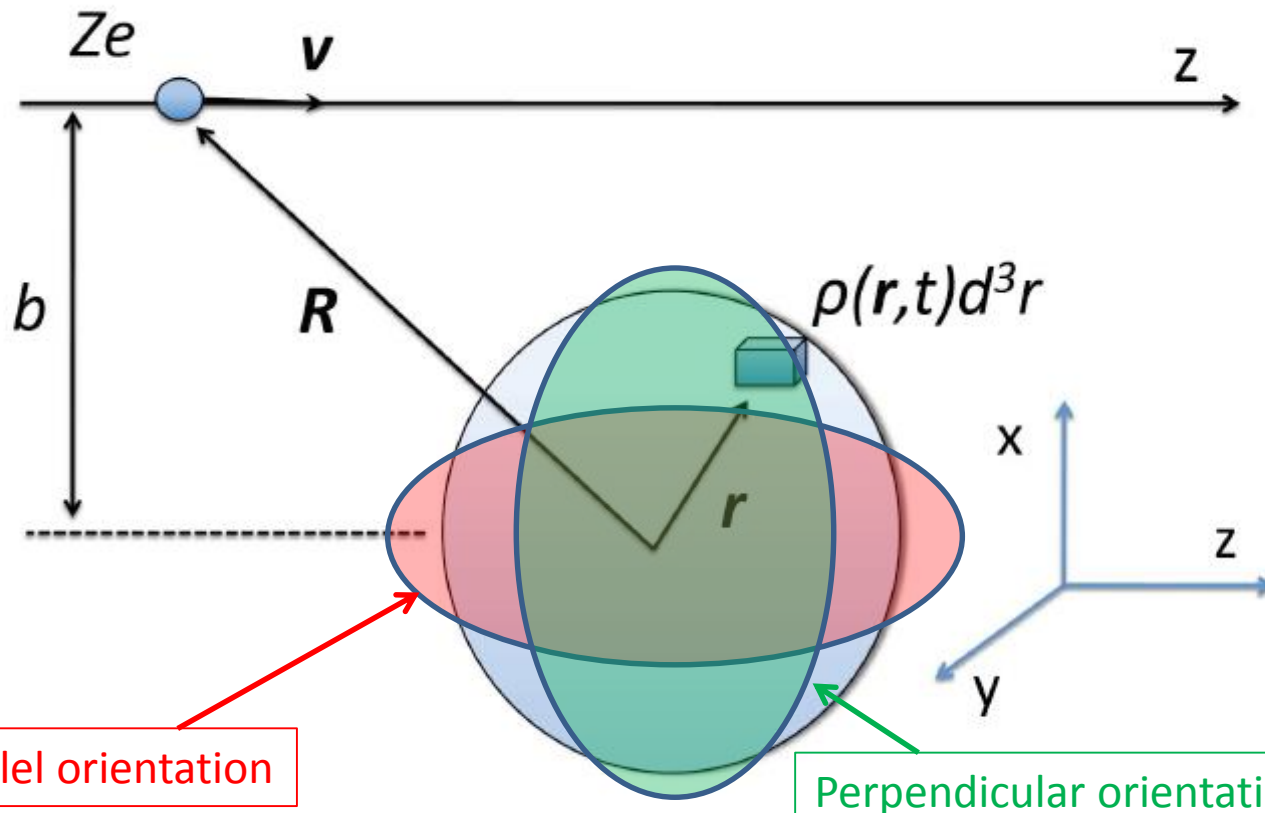
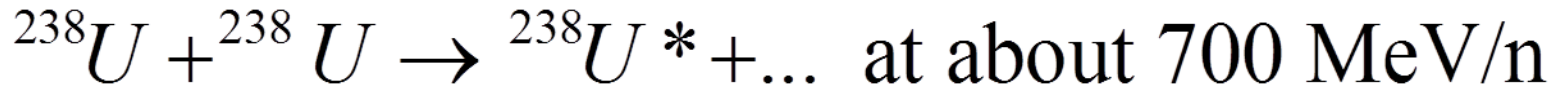
$$S(E) = \sum_{\nu} |\langle \nu | \hat{F} | 0 \rangle|^2 \delta(E - E_{\nu})$$

$$S(\omega) = \text{Im}\{\delta F(\omega)/[\pi f(\omega)]\}$$

(gamma,n) reaction
through the excitation of GDR



Beyond linear regime: *Relativistic Coulomb excitation*



The coordinate transformation has been applied to keep CM in the center of the box at all times.

Coupling to e.m. field:

$$\vec{E} = -\vec{\nabla}\phi - \frac{1}{c} \frac{\partial \vec{A}}{\partial t}$$

$$\vec{B} = \vec{\nabla} \times \vec{A}$$

$$\vec{\nabla}\psi \rightarrow \vec{\nabla}_A\psi = \left(\vec{\nabla} - i\frac{e}{\hbar c}\vec{A}\right)\psi$$

$$\vec{\nabla}\psi^* \rightarrow \vec{\nabla}_{-A}\psi^* = \left(\vec{\nabla} + i\frac{e}{\hbar c}\vec{A}\right)\psi^*$$

$$i\hbar\frac{\partial}{\partial t}\psi \rightarrow \left(i\hbar\frac{\partial}{\partial t} - e\phi\right)\psi$$

which implies that $\vec{\nabla}\psi\psi^* \rightarrow \vec{\nabla}\psi\psi^*$.

Consequently the densities change according to:

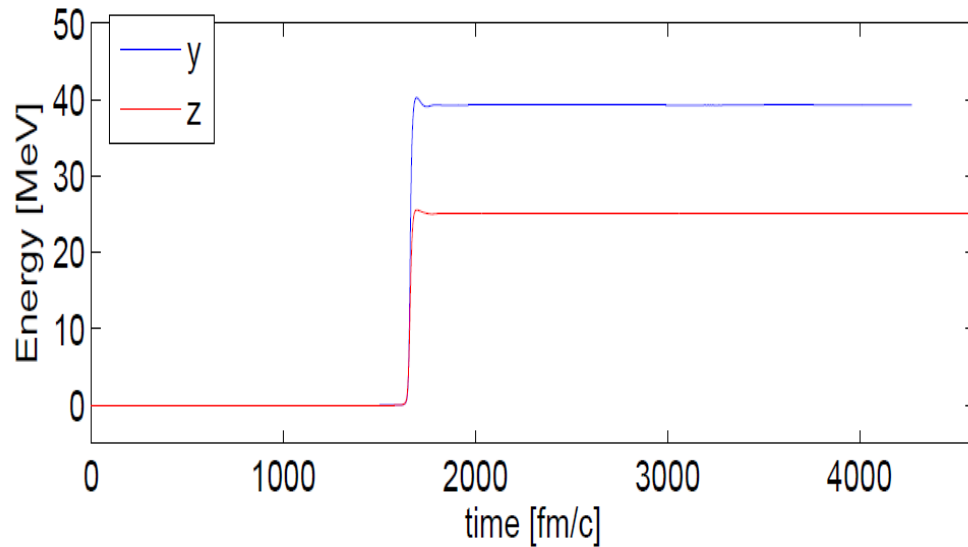
- density: $\rho_A(\mathbf{r}) = \rho_A(\mathbf{r})$
- spin density: $\vec{s}_A(\mathbf{r}) = \vec{s}(\mathbf{r})$
- current: $\vec{j}_A(\mathbf{r}) = \vec{j}(\mathbf{r}) - \frac{e}{\hbar c}\vec{A}\rho(\mathbf{r})$
- spin current (2nd rank tensor): $\mathbf{J}_A(\mathbf{r}) = \mathbf{J}(\mathbf{r}) - \frac{e}{\hbar c}\vec{A} \otimes \vec{s}(\mathbf{r})$
- spin current (vector): $\vec{J}_A(\mathbf{r}) = \vec{J}(\mathbf{r}) - \frac{e}{\hbar c}\vec{A} \times \vec{s}(\mathbf{r})$
- kinetic energy density: $\tau_A(\mathbf{r}) = \left(\vec{\nabla} - i\frac{e}{\hbar c}\vec{A}\right) \cdot \left(\vec{\nabla}' + i\frac{e}{\hbar c}\vec{A}\right) \rho(\mathbf{r}, \mathbf{r}')|_{r=r'}$
 $= \tau(\mathbf{r}) - 2\frac{e}{\hbar c}\vec{A} \cdot \vec{j}(\mathbf{r}) + \frac{e^2}{\hbar^2 c^2}|\vec{A}|^2\rho(\mathbf{r}) = \tau(\mathbf{r}) - 2\frac{e}{\hbar c}\vec{A} \cdot \vec{j}_A(\mathbf{r}) - \frac{e^2}{\hbar^2 c^2}|\vec{A}|^2\rho(\mathbf{r})$
- spin kinetic energy density: $\vec{T}_A(\mathbf{r}) = \left(\vec{\nabla} - i\frac{e}{\hbar c}\vec{A}\right) \cdot \left(\vec{\nabla}' + i\frac{e}{\hbar c}\vec{A}\right) \vec{s}(\mathbf{r}, \mathbf{r}')|_{r=r'}$
 $= \vec{T}(\mathbf{r}) - 2\frac{e}{\hbar c}\vec{A}^T \cdot \mathbf{J}(\mathbf{r}) + \frac{e^2}{\hbar^2 c^2}|\vec{A}|^2\vec{s}(\mathbf{r}) = \vec{T}(\mathbf{r}) - 2\frac{e}{\hbar c}\vec{A}^T \cdot \mathbf{J}_A(\mathbf{r}) - \frac{e^2}{\hbar^2 c^2}|\vec{A}|^2\vec{s}(\mathbf{r})$

Energy deposited for two nuclear orientations (y – perpendicular, z – parallel)

I. Stetcu, C. Bertulani, A. Bulgac, P. Magierski, K.J. Roche
Phys. Rev. Lett. 114, 012701 (2015)

Impact parameter $b=12.2\text{fm}$

Excitation energy (CM motion subtracted)

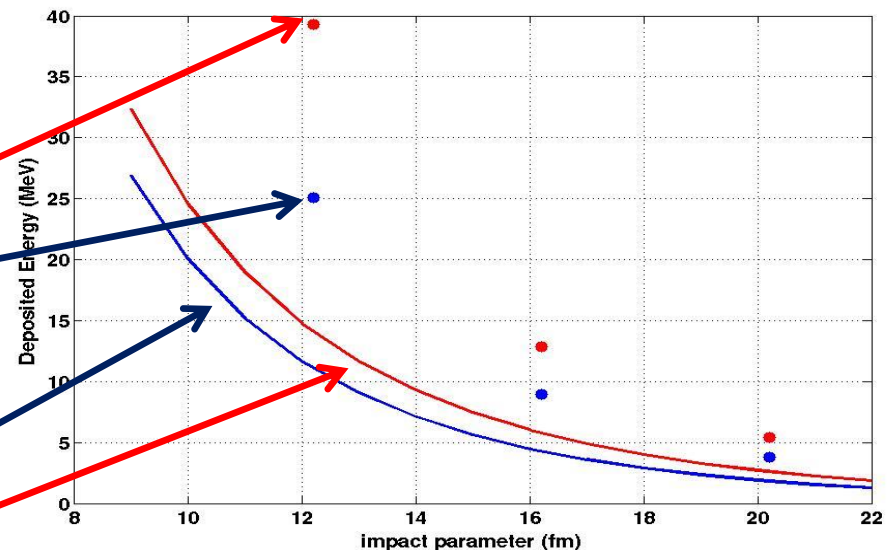


Energy transferred to the target nucleus in the form of internal excitations

TDSLDA – perpendicular orientation

TDSLDA – parallel orientation

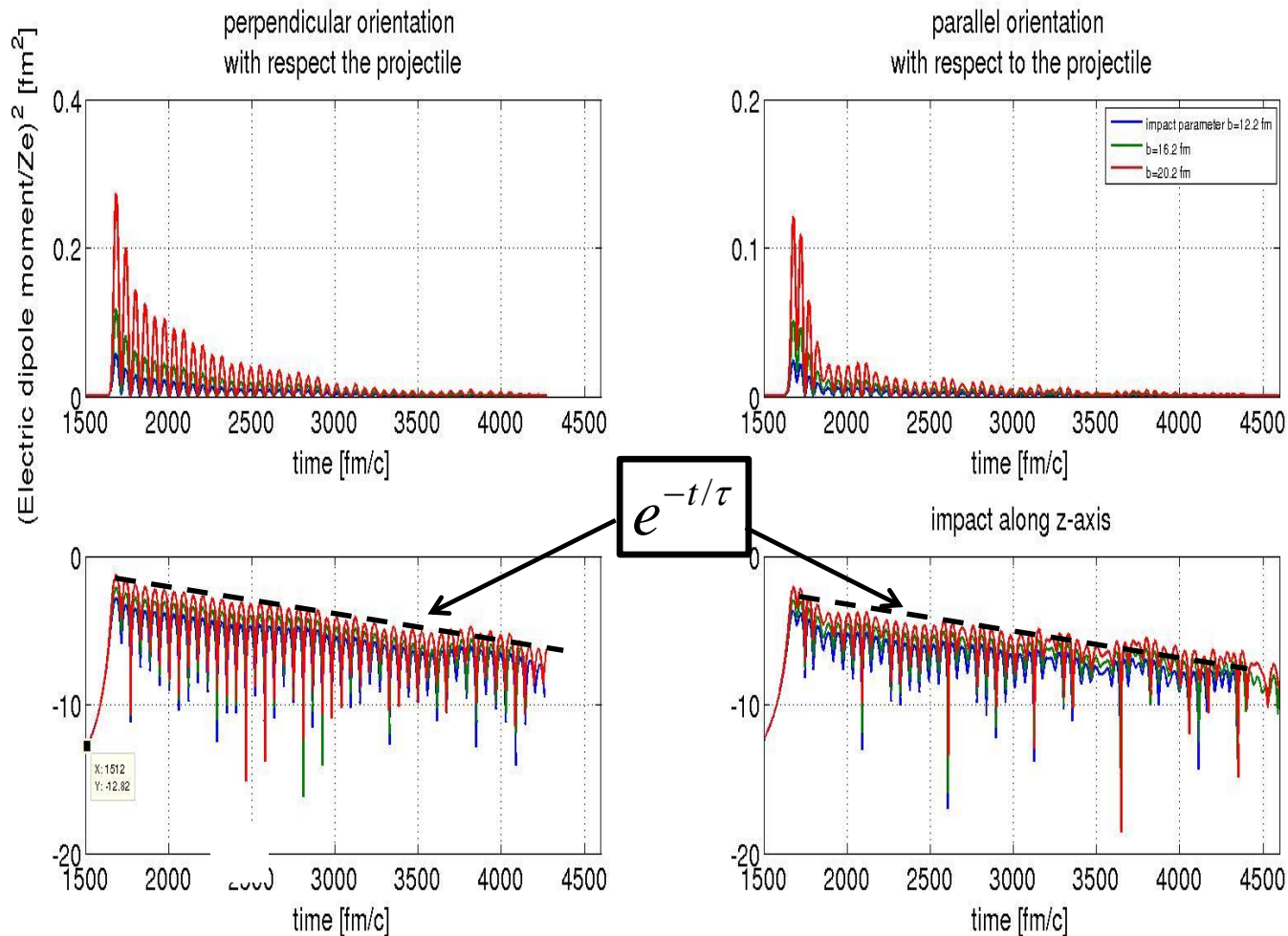
Goldhaber-Teller like model:
proton and neutron density distributions
oscillating against each other



One body dissipation

Let us assume that the collective energy of dipole oscillation is proportional the square of the amplitude of electric dipole moment:

$$E_{coll}(t) \propto [D_{max}(t)]^2$$

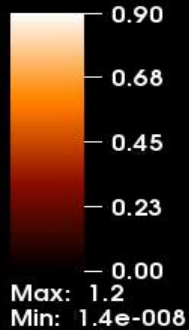


$$E_{coll}(t) \propto e^{-t/\tau}; \quad \tau \approx 500 \text{ fm} / c \Rightarrow \Gamma_{\downarrow} \approx 0.4 \text{ MeV}$$

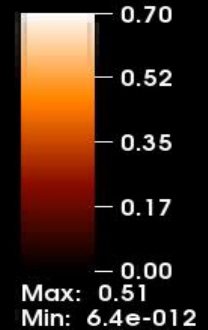
Induced nuclear fission by neutron capture: pairing dynamics

Fission of ^{240}Pu at excitation energy $E_x = 8.08$ MeV

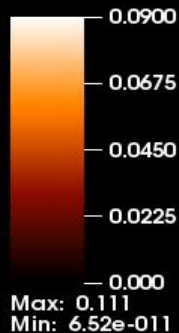
Neutron pairing gap (MeV)



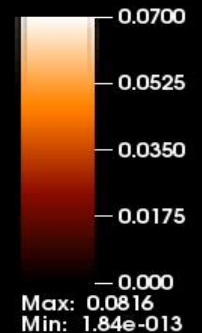
Proton pairing gap (MeV)



Neutron density (fm^{-3})



Proton density (fm^{-3})



Time= 0.000000 fm/c

Bulgac, Magierski, Roche, Stetcu, PRL 116, 122504 (2016)

Induced nuclear fission by neutron capture

Fission of ^{240}Pu at excitation energy $E_x = 8.05; 7.91; 8.08$ MeV

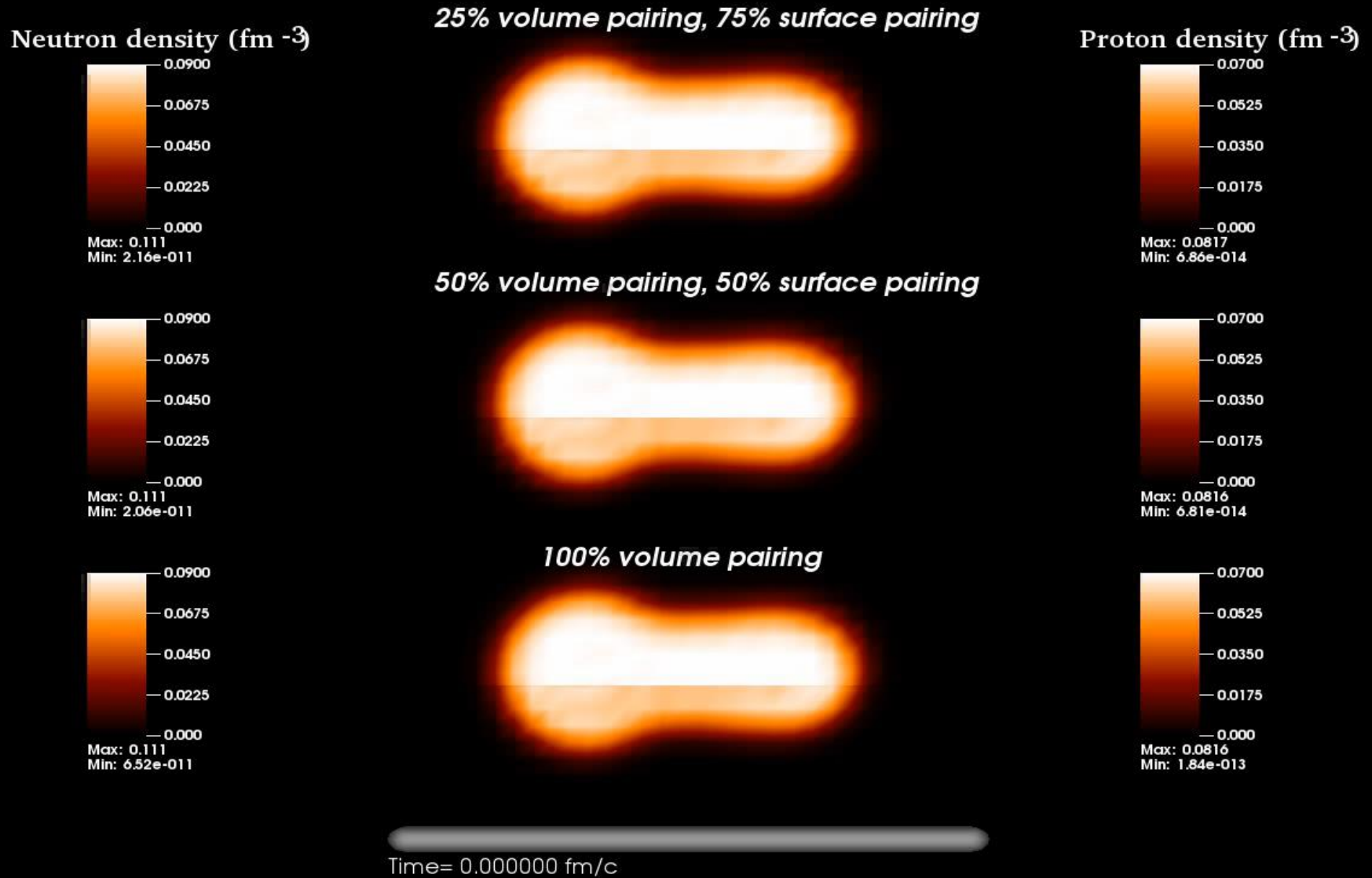


TABLE I: The simulation number, the pairing parameter η , see Eq. (1), the excitation energy (E^*) of the mother ${}^{240}_{94}\text{Pu}_{136}$ and of the daughter nuclei ($E_{H,L}^*$), the equivalent neutron incident energy (E_n), the starting initial quadrupole moment, the “saddle-to-scission” time, the total kinetic energy (TKE), atomic ($A_{H,L}$), neutron ($N_{H,L}$) and proton ($Z_{H,L}$) numbers of the heavy and light fragments, and the number of neutrons (ν), estimated using a Hauser-Feshbach approach and experimental neutron separation energies [8, 68, 69]. Units are MeV, fm^2 and fm/c where appropriate.

S#	η	E^*	E_n	Q_{zz}	S-S time	TKE	A_H	A_L	N_H	N_L	Z_H	Z_L	E_H^*	E_L^*	ν_H	ν_L
S1	0.75	8.05	1.52	16,500	14,419	182	136.0	104.0	83.2	62.8	52.8	41.2	5.26	17.78	0	1.9
S2	0.5	7.91	1.38	16,500	4,360	183	133.7	106.3	82.0	64.0	51.7	42.3	9.94	11.57	1	1
S3	0	8.08	1.55	16,500	14,010	180	134.5	105.5	82.4	63.6	52.1	41.9	3.35	29.73	0	2.9
S4	0	6.17	-0.36	19,000	12,751	181	136.1	103.9	83.4	62.6	52.7	41.3	7.85	9.59	1	1

$$\text{TKE} = 177.80 - 0.3489E_n \quad [\text{in MeV}],$$

Nuclear data evaluation, Madland (2006)

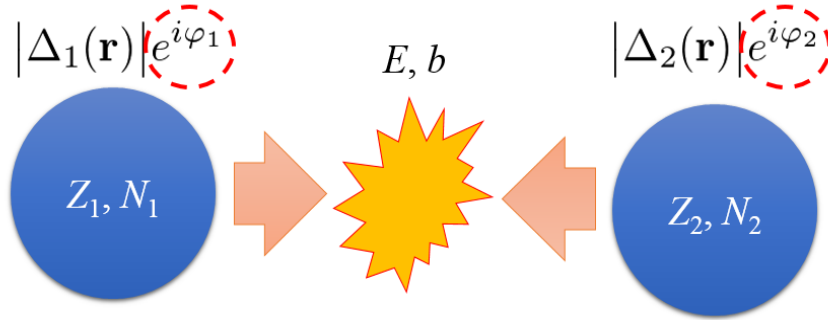
estimated TKEs slightly overestimate the observed values by no more than 3 - 6 MeV.

This is indicative of the fact that in our simulations the system scissions a bit too early.

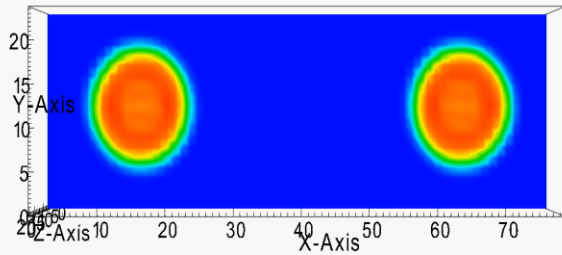
The evaluated average number of emitted neutrons in this case is close to 3, which is higher than the values we estimate.

Solitonic excitations in nuclear reactions

$$\Delta\varphi (= \varphi_2 - \varphi_1)$$



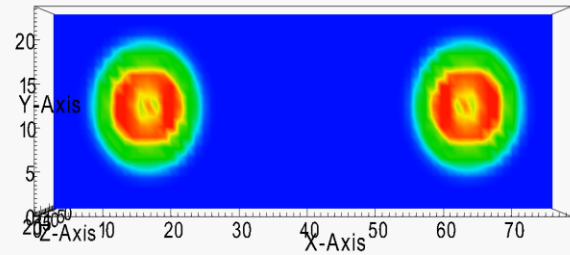
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 Cycle: 0 Time: 993.648
 Var: density
 Units: fm⁻³
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 0.1200
 0.08000
 0.04000
 0.000
 Max: 0.1529
 Min: 1.455e-10



$$\Delta\varphi = \pi$$

$$\Delta\varphi = 0$$

Pseudocolor
 Dis: colleson12_delta_n.0000
 Cycle: 0 Time: 993.648
 Var: delta_obs
 Units: MeV
 1.700
 1.275
 0.8500
 0.4250
 0.000
 Max: 1.790
 Min: 1.832e-07

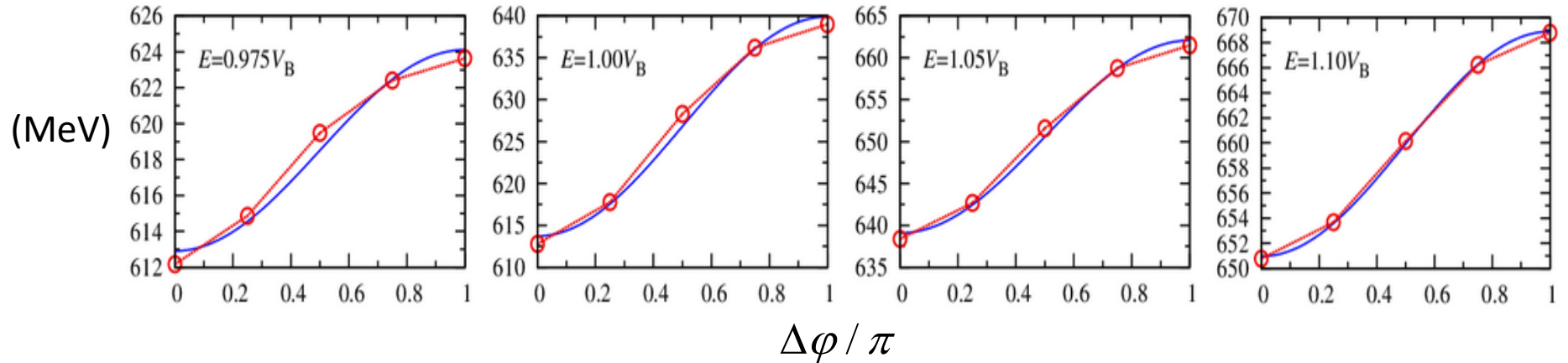


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 Wed Apr 20 18:57:21 2016



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 Wed Apr 20 18:58:55 2016

$$E_{\text{kin}}(E, \Delta\varphi) = A(E) - B(E) \cos(\Delta\varphi)$$



From the talk of K. Sekizawa

Energy of the junction between two superfluids of different phase:

$$E \propto |\Delta|^2 \sin^2 \left(\frac{\Delta\varphi}{2} \right)$$

Summary:

- TDSLDA offers insights into nuclear processes which are either not easy or impossible to obtain in the laboratory.
- We were able to describe nucleus-vortex interaction from the dynamics without any symmetry constraints, taking into account internal degrees freedom of the nucleus and the vortex.
- We extracted the force per unit of the vortex length, which can be used as an input for simplified large scale calculations for the neutron star crust (eg. based on filament model).
- For neutron matter we were able to extract the effective mass of impurity (nucleus) immersed in the neutron environment.
- The quality of the agreement with experimental observations for fission is surprisingly good, taking into account we made no effort to reproduce any measured data.
- TDSLDA predicts much longer time-scales for fission, even though it takes into account one-body dissipation only (both window and wall mechanisms are present). The nuclear system superficially behaves like an extremely viscous system, but the collective motion at the same time is not overdamped.

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